

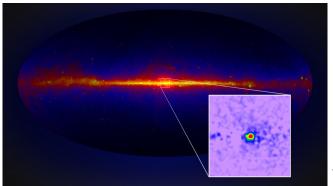
Neutron Stars in the Dense Stellar Cluster at the Center of the Galaxy

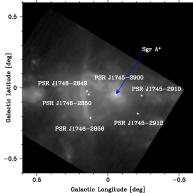
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Collaborators: Jonathon Kocz (Caltech), Shinji Horiuchi (CSIRO)

Program: Strategic Initiative R&TD







PROJECT OBJECTIVE

- Survey of the inner region of the galaxy for pulsars with new DSN capability at K-band
- Develop search methodologies and efficient pipelines to detect isolated and binary pulsars in the GC, including Crab-like pulsars with bright single pulse emitters
- Tune up pipelines with observations of radio magnetars, e.g., the GC magnetar PSR J1745-2900

BACKGROUND

- NASA's Fermi LAT instrument has identified excess gamma-ray emission toward the inner part of the GC
- Leading explanations include the possibilities of direct annihilation of DM or a large population of MSPs
- Detection of MSPs in dense nuclear star clusters pose a special challenge for surveys, requiring high frequency, large antennas, wide bandwidths, long integration times

BENEFITS TO NASA & JPL

- Major contribution to Fermi science in tackling an outstanding mystery in high energy astrophysics
- Making use of the new K-band receiving system at the DSN 70-m antenna, which is both a unique JPL asset and the most sensitive centimeter-wavelength antenna in the Southern Hemisphere

RELEVANT PUBLICATIONS

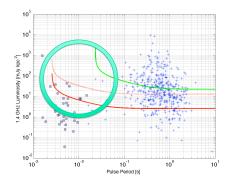
- Majid, W., Pearlman, A., Prince, T., et al., "Postoutburst Radio Observations of the High Magnetic Field Pulsar PSR J1119-6127", ApJL, 834, L2 (2017).
- Pearlman, A., Majid, W., Prince, T., et al., "Pulse Morphology of the Galactic Center Magnetar PSR J1745-2900", ApJ, 866, 160 (2018).
- Pearlman, A., Majid W., & Prince "Observations of Radio Magnetars with the Deep Space Network", Advances in Astronomy, vol 2019, 6325183 (2019).

National Aeronautics and Space Administration

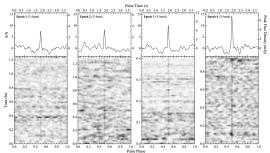
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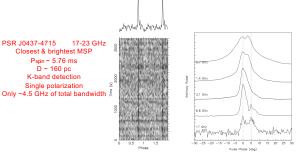
Detectability of pulsars in the GC region



Average pulse profiles of the GC Magnetar J1745-2900



Highest frequency detection of MSP PSR J0437-4715



Previous highest frequency detection Keith+2011, MNRAS